

## Asteroid families spin/shape models database to be supported by ProjectSoft robotic observatory



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**Abstract:** In our recent work (Hanuš et al. 2013, A&A 559, A134) we studied dynamics of asteroid families constrained by the distribution of pole latitudes vs semimajor axis. The model contained the following ingredients: (i) the Yarkovsky semimajor-axis drift, (ii) secular spin evolution due to the YORP effect, (iii) collisional reorientations, (iv) a simple treatment of spin-orbit resonances and (v) of mass shedding. The major limitation of our model stems from the fact that the number of observed family members with known pole latitudes is limited.

We thus outline an ongoing construction of the ProjectSoft robotic observatory called "Blue Eye 600", which will support our efforts to complete the sample of shapes for a substantial fraction of (large) family members. Dense photometry will be targeted in such a way to maximize a possibility to derive a new pole/shape model. Other possible applications of the observatory include: (i) fast resolved observations of fireballs (thanks to a fast-motion capability, up to 90 degrees/second), or (ii) an automatic survey of a particular population of objects (MBAs, NEAs, variable stars, novae etc.).

We also suggest to use a different (complementary) approach, based on distribution functions of shape parameters. Based on >1000 old and new convex-hull shape models, we will be able to construct the distributions of suitable quantities (ellipticity, normalized facet areas, etc.) and we shall discuss a significance of differences among asteroid populations. A check for outlier points may then serve as a possible identification of (large) interlopers among "real" family members. This has also implications for SPH models of asteroid disruptions which can be possibly further constrained by the shape models of resulting fragments. Up to now, the observed size-frequency distribution and velocity field were used as constraints, sometimes allowing for a removal of interlopers (Michel et al. 2011, Icarus 211, 535).



**DAMIT**: Database of Asteroid Models from Inversion Techniques (http://astro.troja.mff.cuni.cz/ projects/asteroids3D) now contains models for **381** asteroids. Most of the models were derived from disk-integrated photometry by the lightcurve inversion, some of them were scaled or refined by adaptive optics disk-resolved images, thermal infrared data, or data from stellar occultations. We plan to use DAMIT for models obtained in the framework of the distributed computing project Asteroids@home (http://asteroidsathome.net).



2.00 2.05 2.10 2.15 2.20 2.25 2.30 2.35 2.40

**ProjectSoft:** The company specializes in renovation of existing astronomical telescopes, domes, spectrographs and other astronomical instruments, requiring computer control, robotic and automatic operations. We are able to guarantee quick implementation including all turn-key deliveries. We use only commercial off-the-shelf solutions and components which significantly increase lifetime of the modernized system. Moreover, this solves the essential problem of astronomical devices which is the availability of spare parts and maintenance. This way they are world-wide available regardless of the supplier.

 $\rightarrow$  we use this experience to construct a fully robotized observatory **BlueEye 600** 

(an automated planner). Right column: a small part of the **aitel** source code written in Python









Figure 1: An example of the Flora family spin-orbital evolution; (a) the observed proper semimajor axis  $a_p$  vs the absolute magnitude H; (b)  $a_{p}$  vs pole latitude  $\beta$ , with sizes D denotes by symbols; (c) the same for a simulation of large asteroids (D > 30 km); and (d) small asteroids. Finally, (e) a sort of  $\chi^2$  metric vs time t.

Figure 3: An ongoing completion of the alt-azimuth mount (to be equipped with Officina Stellare RiLA 600 f/5 reflector, MI CCD camera with an E2V chip 42-20-BI, >90% QE, FOV 0.52 deg). Most of the electrical equipment (including UPS power supply) is located in a separate container. The place at the Ondřejov observatory is ready.



Acknowledgements: This project was supported by Technology Agency of the Czech Republic (TA ČR), grant no. TA03011171 "Development of the technologies for fast robotic observatories and laser communicating systems".

